

Ngcer Calibration Update

I have been working on the noble gas calibration.

I looked at two methods,

1. The Poisson fit that that Cameroon Cotton developed
2. The multi-Gaussian fit that was used to calibrate the aerogel.

I have concluded that the Poisson method is much easier to get working properly, and has given me one set of parameters for both 2021, and 2022.

Poisson Method

Uses a modified Poisson function: $f(x) = A \frac{\lambda^{\frac{x}{\mu}} e^{-\frac{\lambda}{\mu}}}{\Gamma(\frac{x}{\mu} + 1)}$

- Where the λ/μ term is the calibration parameter of interest
- Cameron has a full write up, including the derivation here:
 - https://github.com/heinricn/hallc_replay_lt/blob/LTSep_Analysis_2022/CALIBRATION/shms_ngcer_calib/SHMS_NGCER_CALIBRATION.pdf
- I left all the stuff he wrote about here the same, only changing some of the cut values and fit windows.

Gaussian Method

Background

This method cites the paper: [https://doi.org/10.1016/0168-9002\(94\)90183-X](https://doi.org/10.1016/0168-9002(94)90183-X)

The function as it is written in the paper is: $S_{\text{real}} = B + S$

$$S_{\text{real}}(x) \approx \left\{ \frac{(1-w)}{\sigma_0 \sqrt{2\pi}} \exp\left(-\frac{(x-Q_0)^2}{2\sigma_0^2}\right) + w\theta(x-Q_0) \right. \\ \left. \times \alpha \exp[-\alpha(x-Q_0)] \right\} e^{-\mu} + \sum_{n=1}^{\infty} \frac{\mu^n e^{-\mu}}{n!} \\ \times \frac{1}{\sigma_1 \sqrt{2\pi n}} \\ \times \exp\left(-\frac{(x-Q_0-Q_{\text{sh}}-nQ_1)^2}{2n\sigma_1^2}\right), \quad (10)$$

Spectrum

This is normalized to 1, so I added 2 additional parameters to make: $f = c1*B + c2*S$

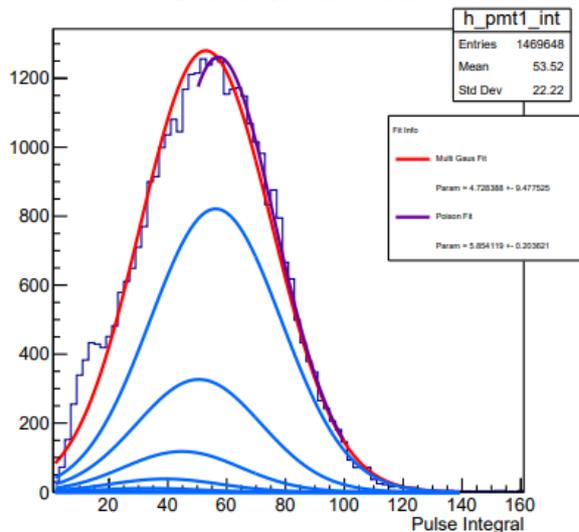
Root did not like my implementation of this function as a for loop (would not fit at all), so I had to hard code the function up to the desired number of terms.

I used 10 Gaussians in what I'm showing, although I tried 15 and that had little effect, but 7 is too few.

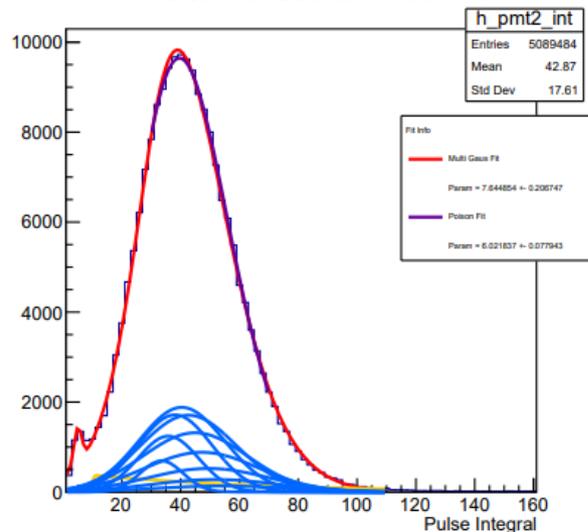
Calibration

- I calibrated 6 π^- settings from across 2021 and 2022, combining the entire setting to get good stats.
- The Gaussian fit often failed, perhaps it is possible to get them to fit properly by fiddling with the parameter limits, and initial values, but I am satisfied with the performance of the Poisson method.

PMT 1 Cerenkov Calibration Poisson Fit



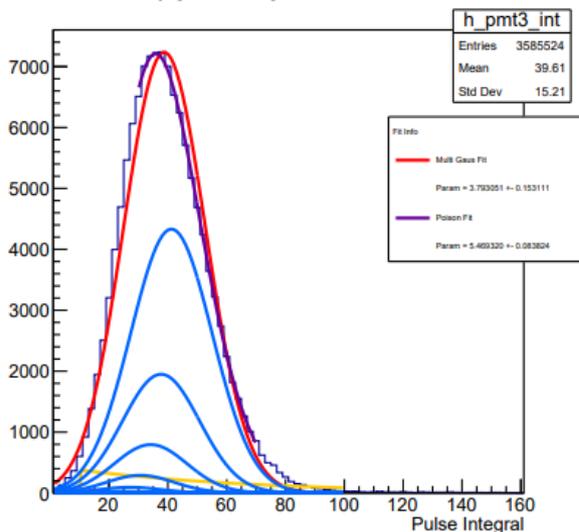
PMT 2 Cerenkov Calibration Poisson Fit



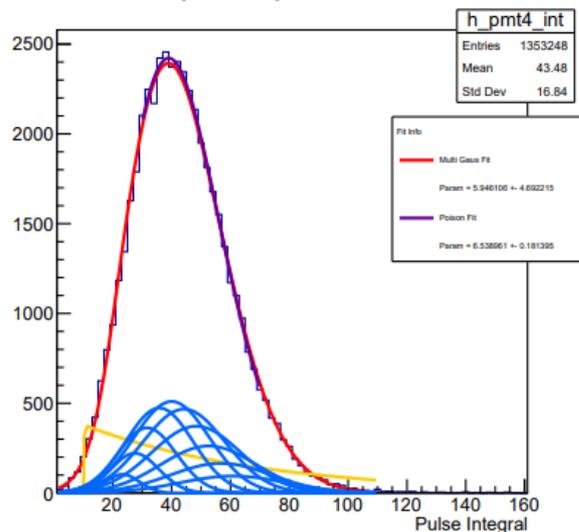
12052

This Run had the fix to the ngcer happen part of the way through the run, so PMTs 1 and 2 have untrustworthy spectrums

PMT 3 Cerenkov Calibration Poisson Fit



PMT 4 Cerenkov Calibration Poisson Fit

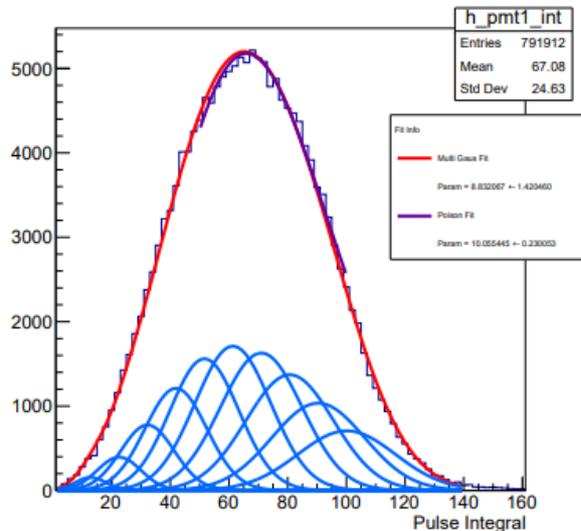


Also, Red line is multi-Gaussian fit

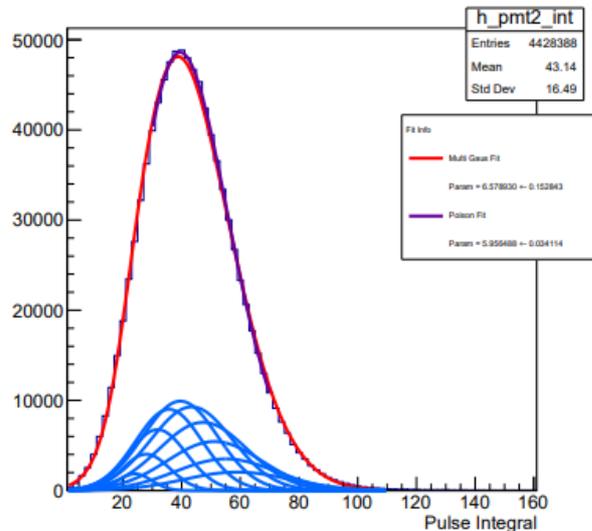
Purple line is the Poisson fit

Blue and yellow are the spectrum and background for the multi-Gauss fit (for debugging)

PMT 1 Cerenkov Calibration Poisson Fit

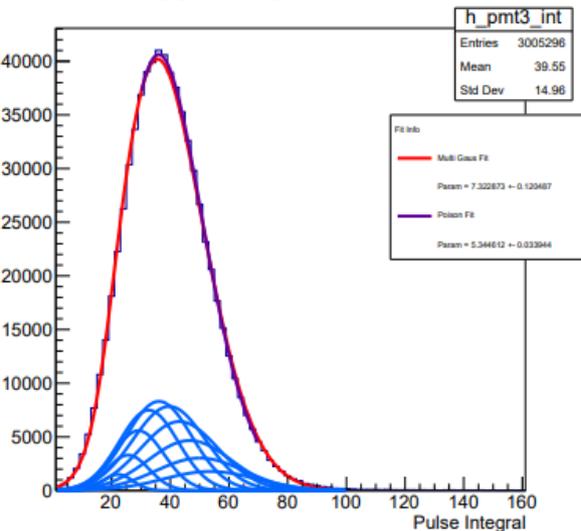


PMT 2 Cerenkov Calibration Poisson Fit

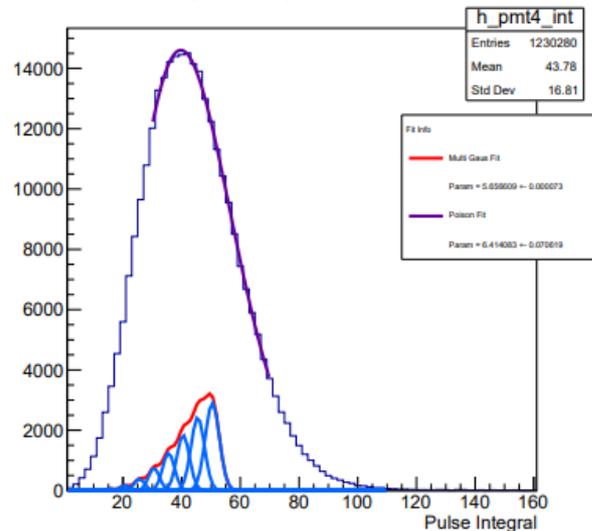


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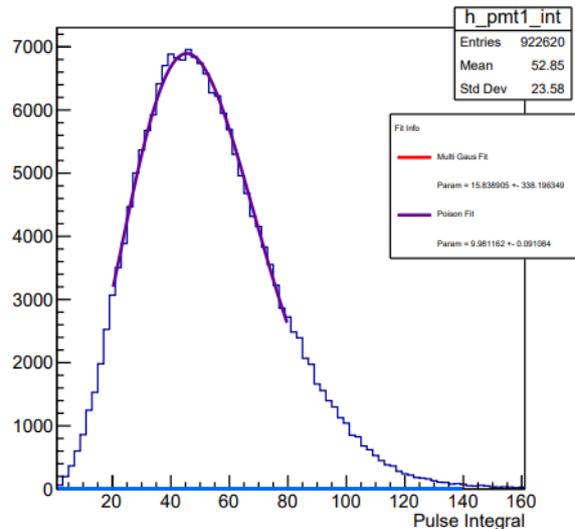
PMT 3 Cerenkov Calibration Poisson Fit



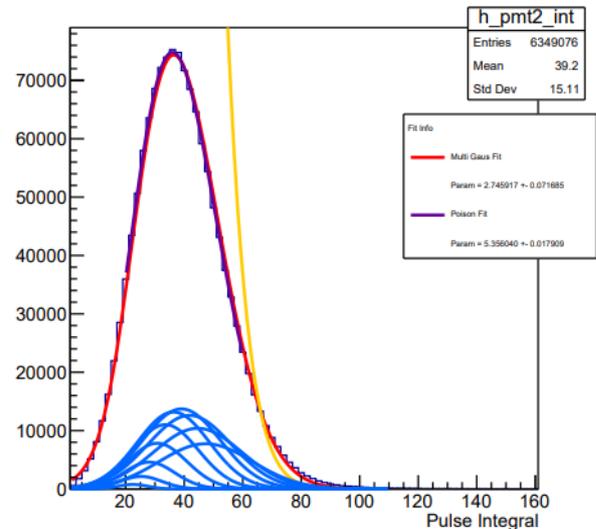
PMT 4 Cerenkov Calibration Poisson Fit



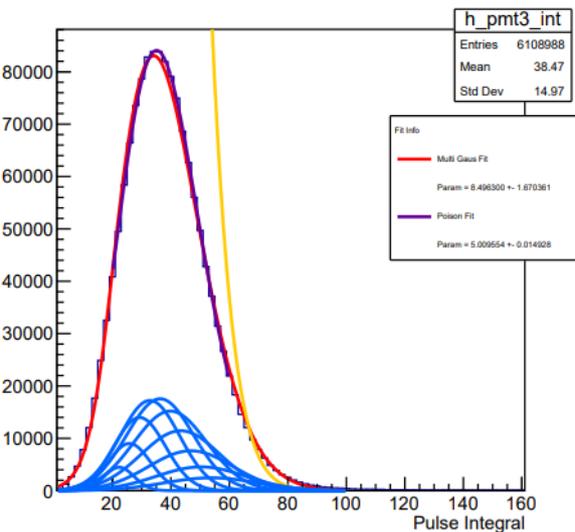
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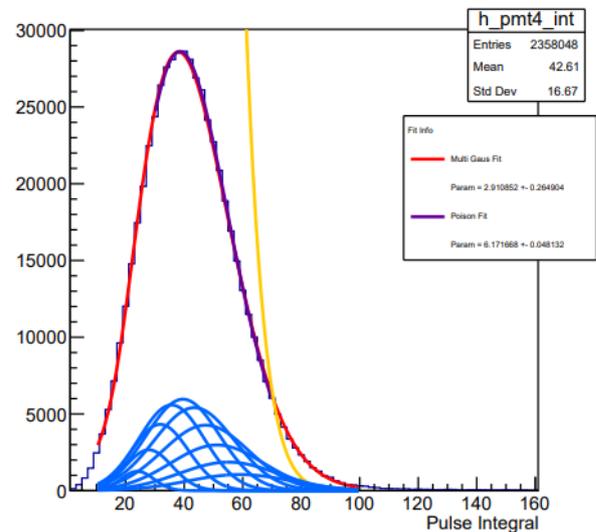
PMT 2 Cerenkov Calibration Poisson Fit



PMT 3 Cerenkov Calibration Poisson Fit

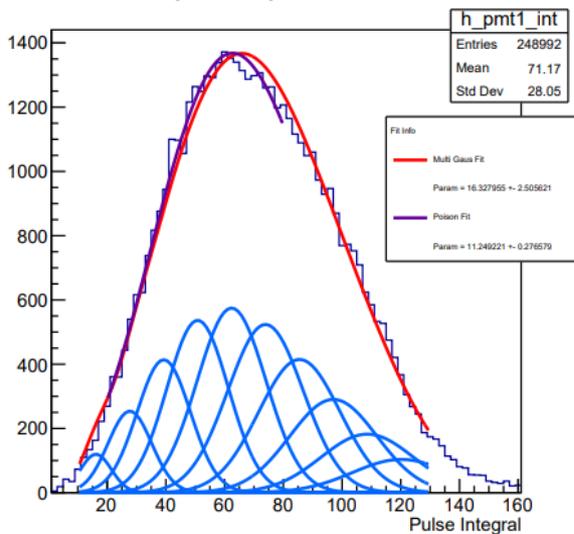


PMT 4 Cerenkov Calibration Poisson Fit

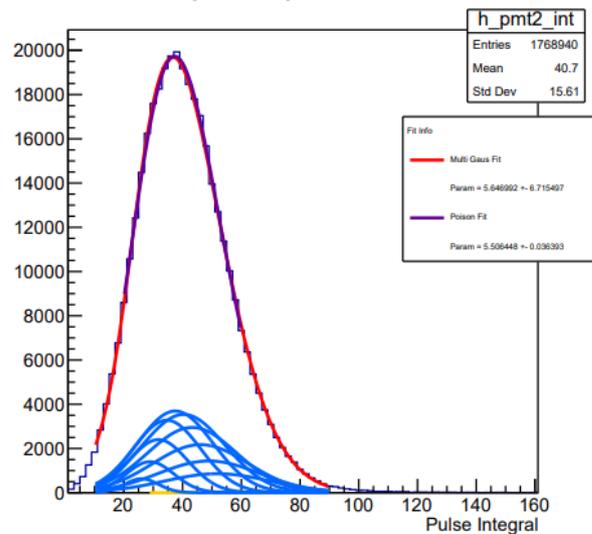


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PMT 1 Cerenkov Calibration Poisson Fit

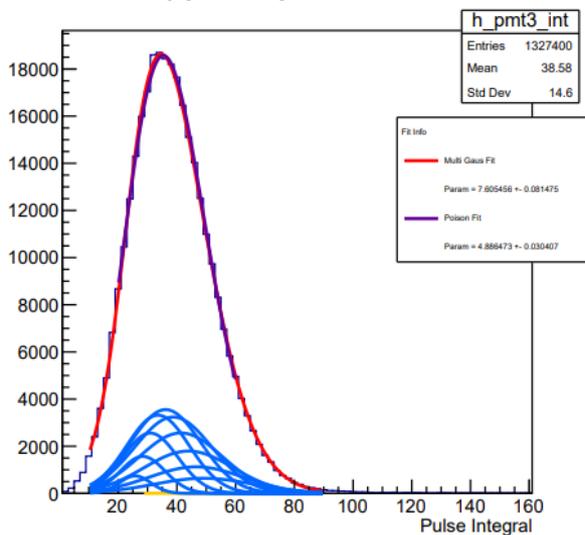


PMT 2 Cerenkov Calibration Poisson Fit

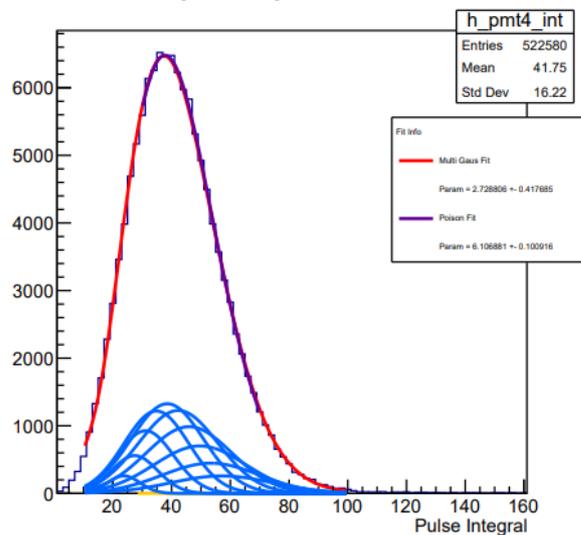


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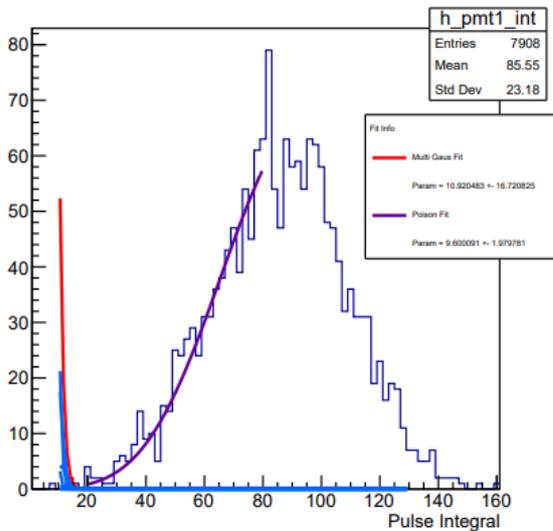


PMT 4 Cerenkov Calibration Poisson Fit

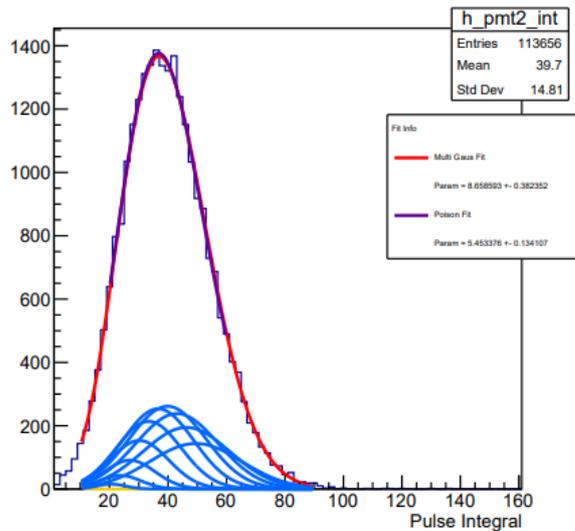


16005

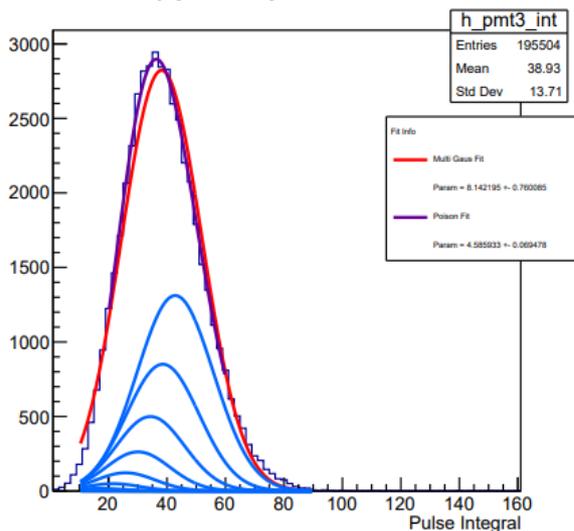
PMT 1 Cerenkov Calibration Poisson Fit



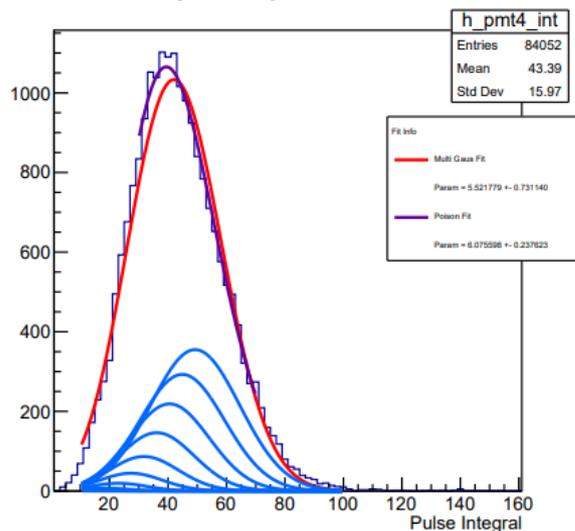
PMT 2 Cerenkov Calibration Poisson Fit



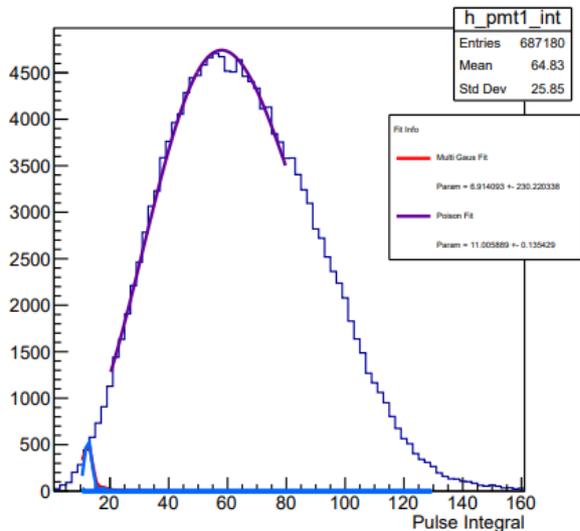
PMT 3 Cerenkov Calibration Poisson Fit



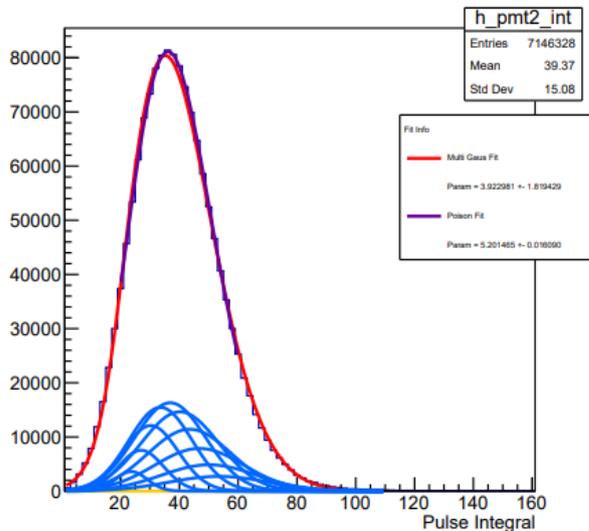
PMT 4 Cerenkov Calibration Poisson Fit



PMT 1 Cerenkov Calibration Poisson Fit

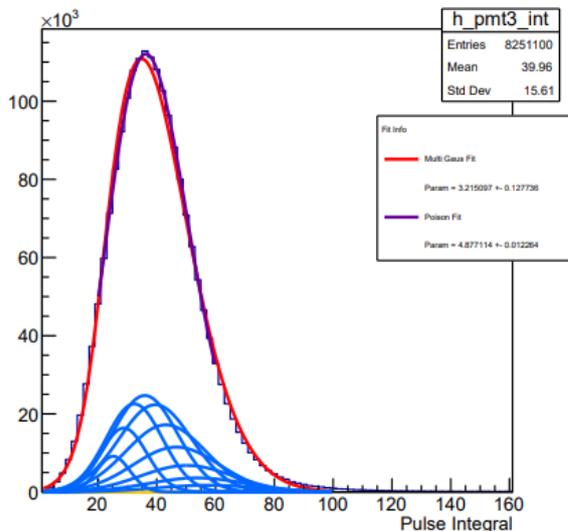


PMT 2 Cerenkov Calibration Poisson Fit

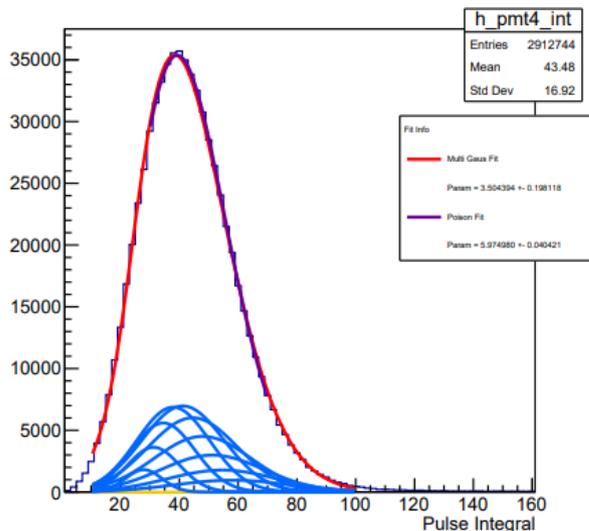


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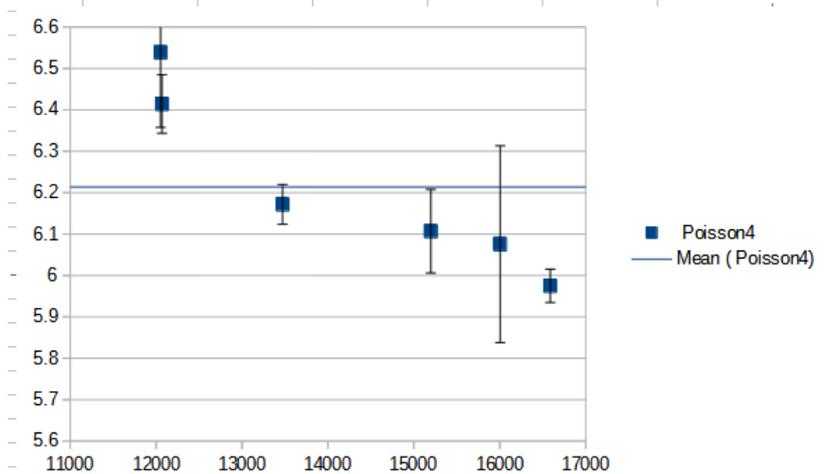
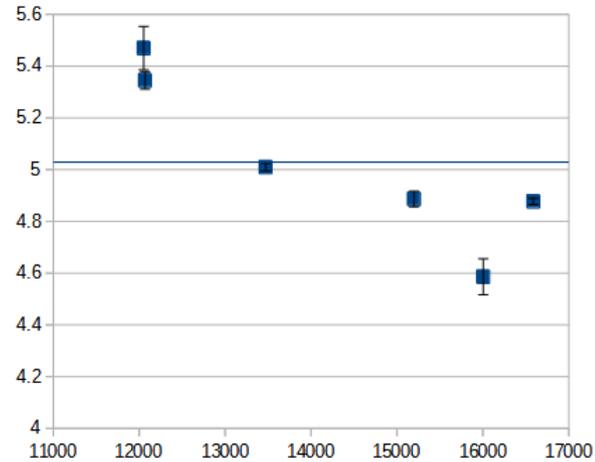
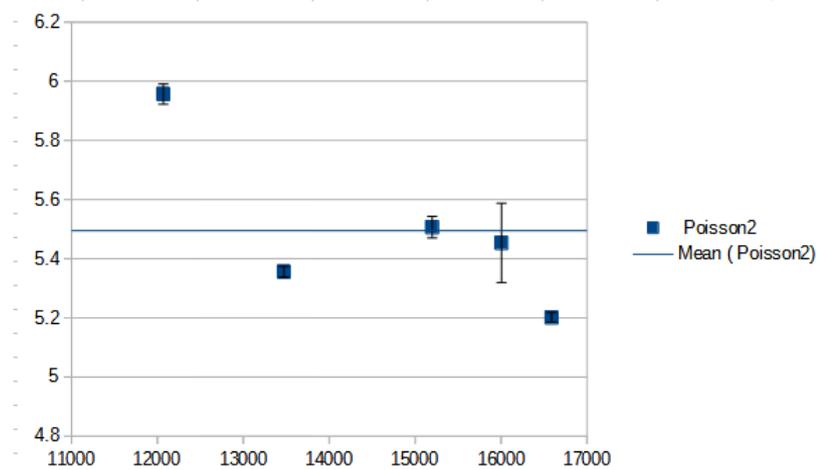
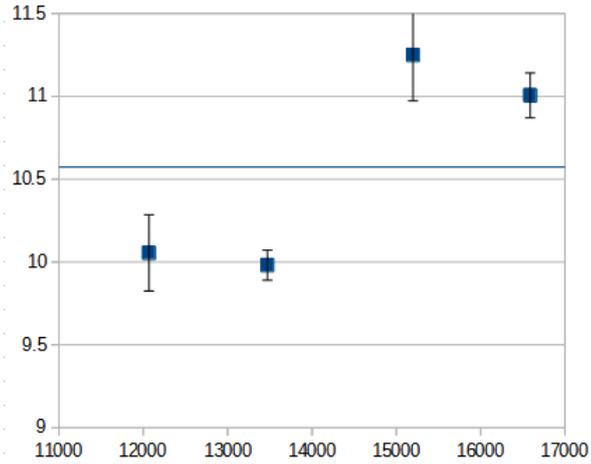
PMT 3 Cerenkov Calibration Poisson Fit



PMT 4 Cerenkov Calibration Poisson Fit



Fit Overview



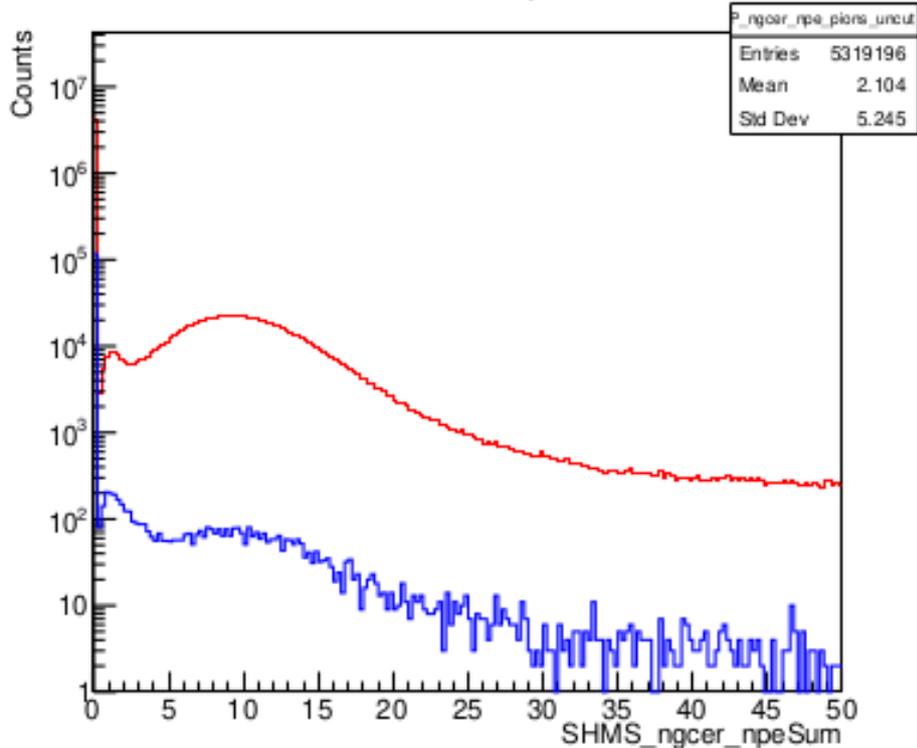
Calibration Conclusions

- I excluded the values from pmts 1 and 2 from 12052, as well as from pmt 1 in 16005. then took the average of the rest.
- The parameters look very stable, so I don't think any further investigation is needed.
- Next is comparison with online, which we expect to be very different as we had a very old set of parameters while we were running.

Q^2 1.60 W 3.08 center high ε LD-

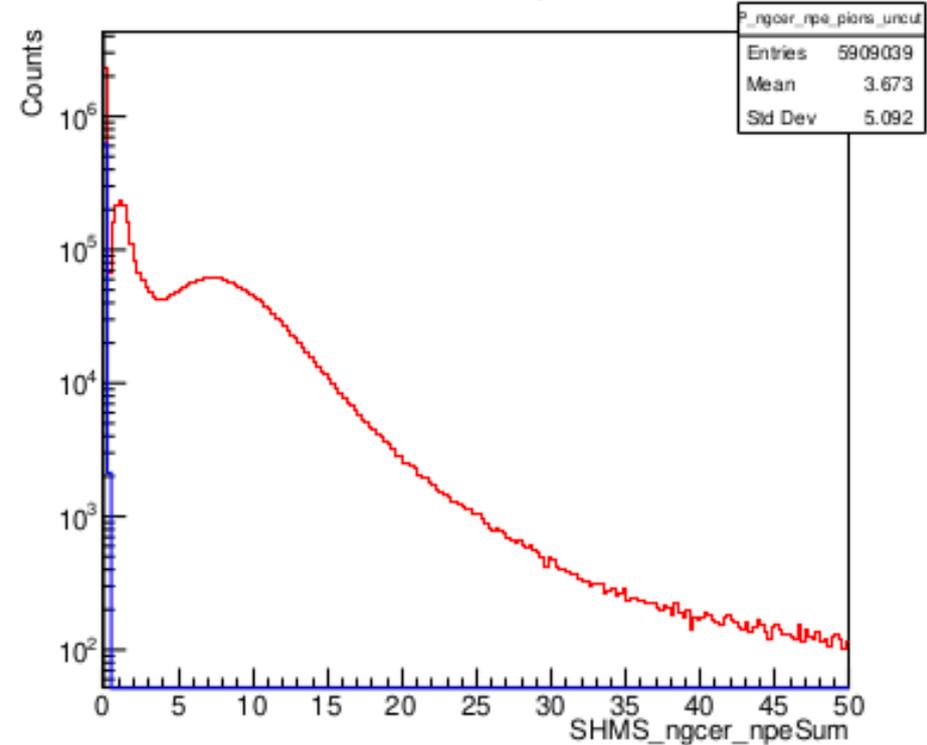
Offline

SHMS NGC npeSum



Online

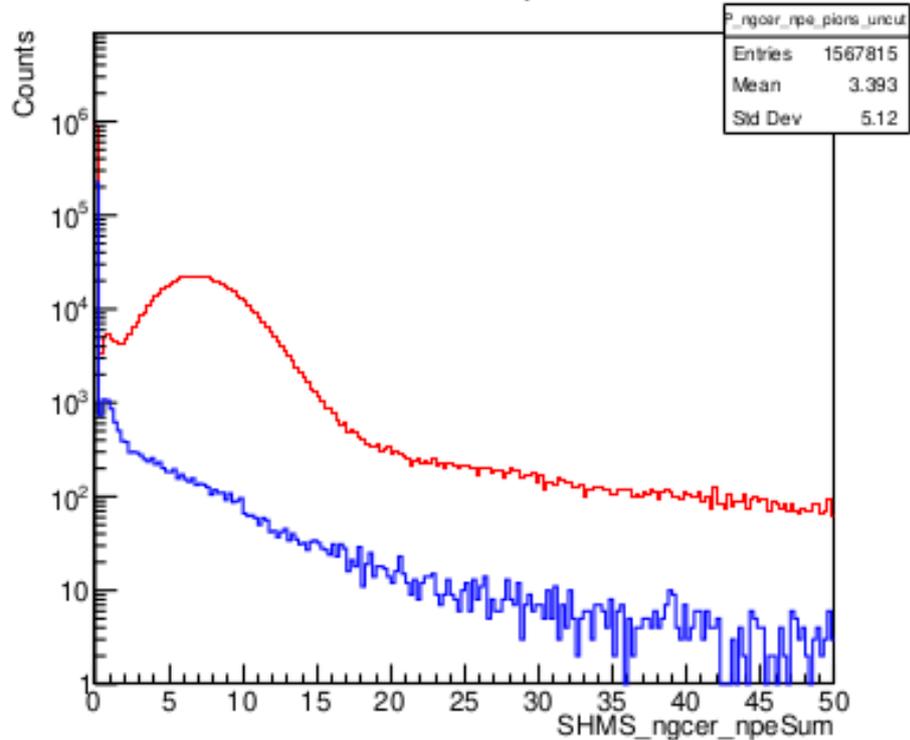
SHMS NGC npeSum



Q^2 3.85 W 2.62 center high ε LD-

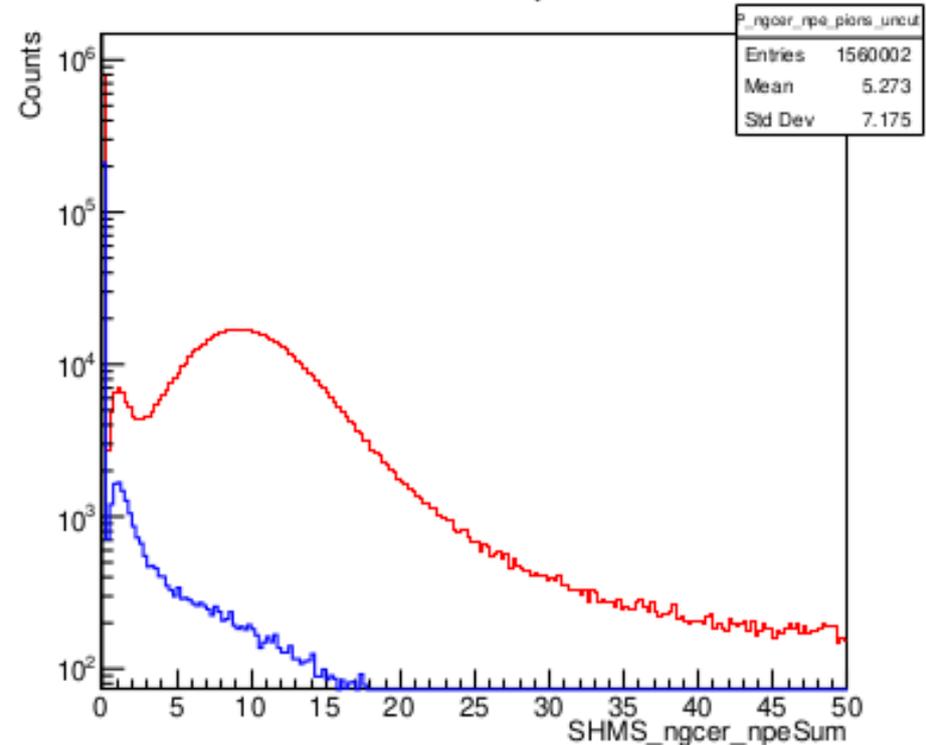
Offline

SHMS NGC npeSum



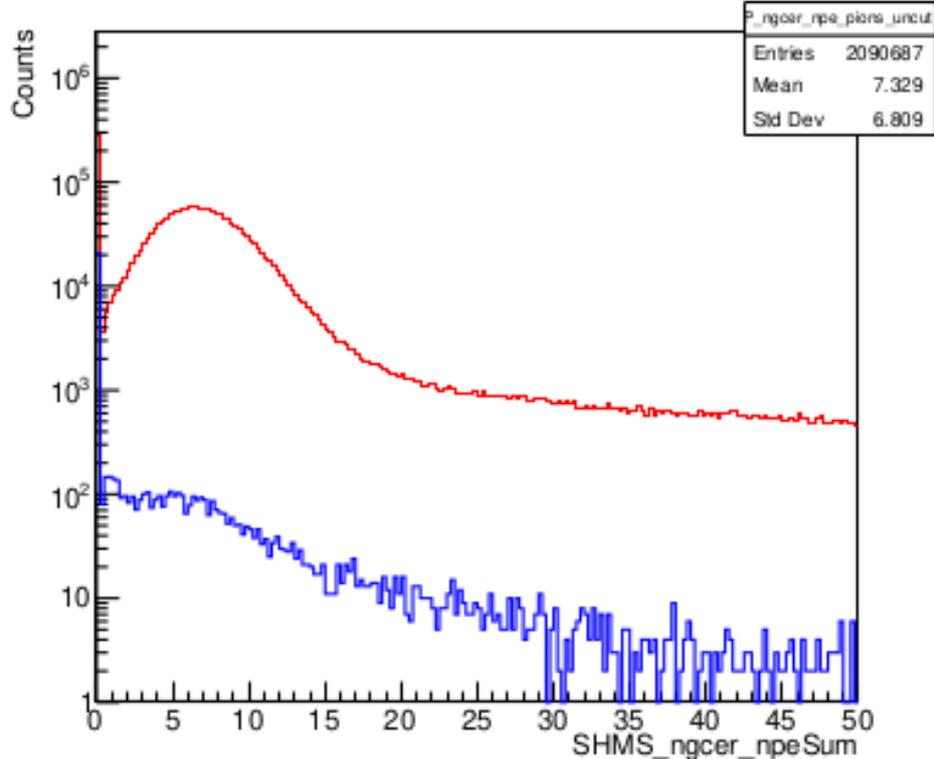
Online

SHMS NGC npeSum

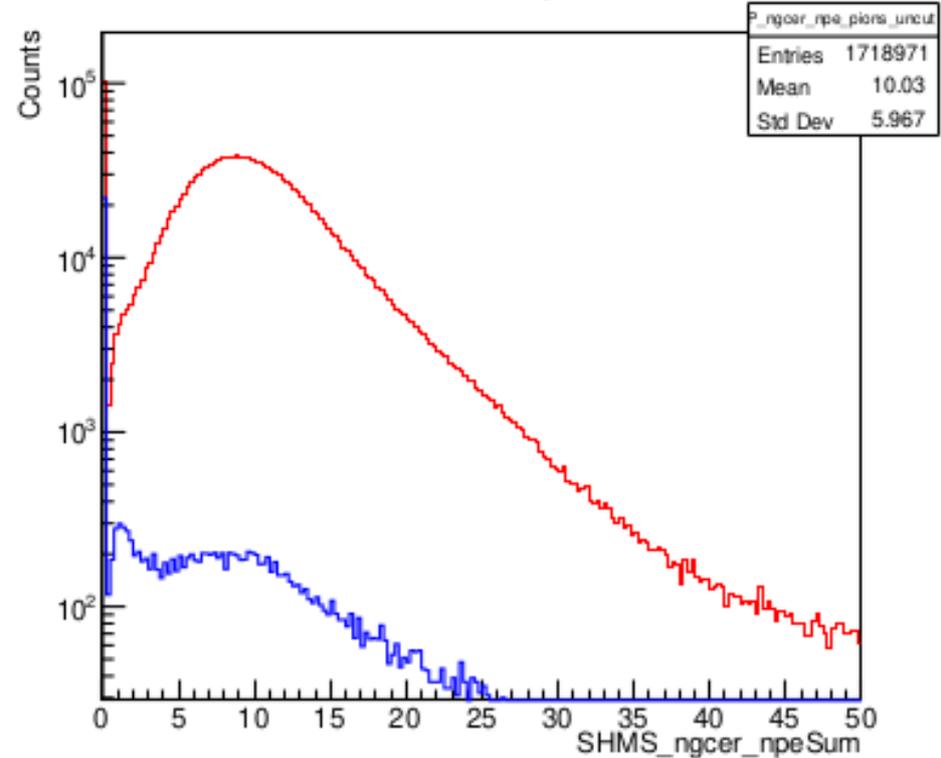


Q² 3.85 W 3.07 center low ϵ LD-

Offline
SHMS NGC npeSum

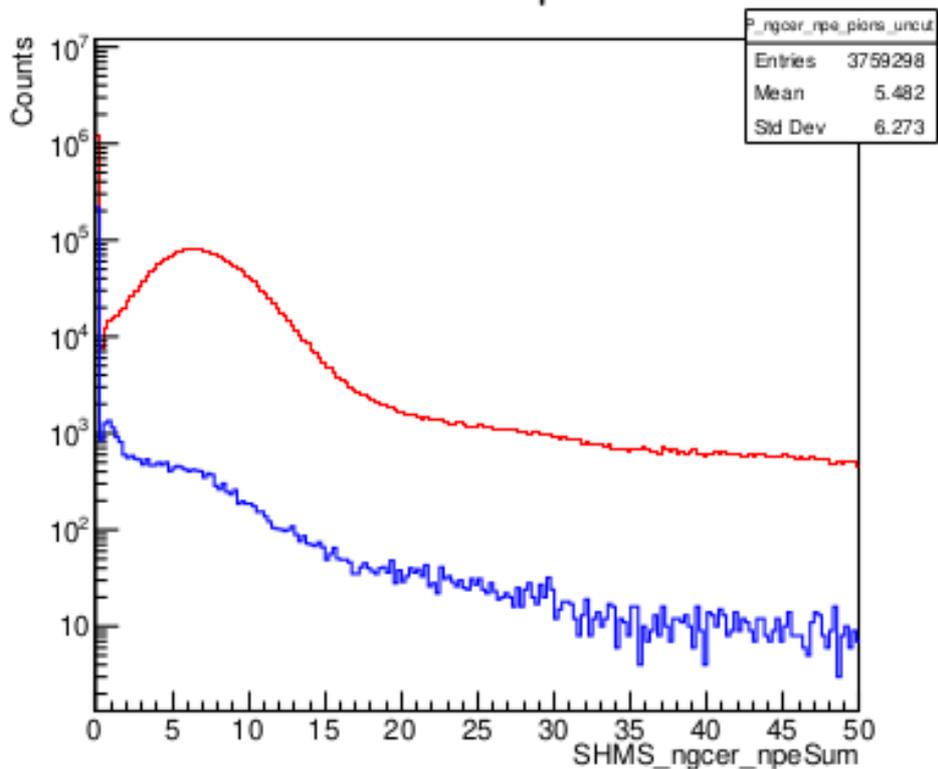


Online
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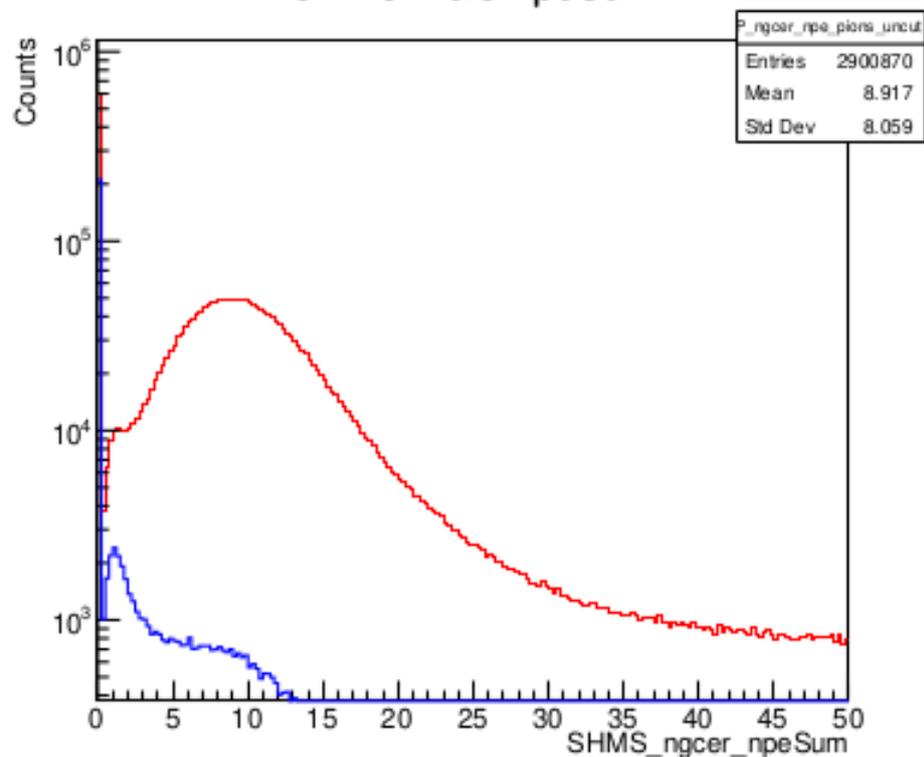


Q^2 3.85 W 3.07 Right high ε LD-

Offline
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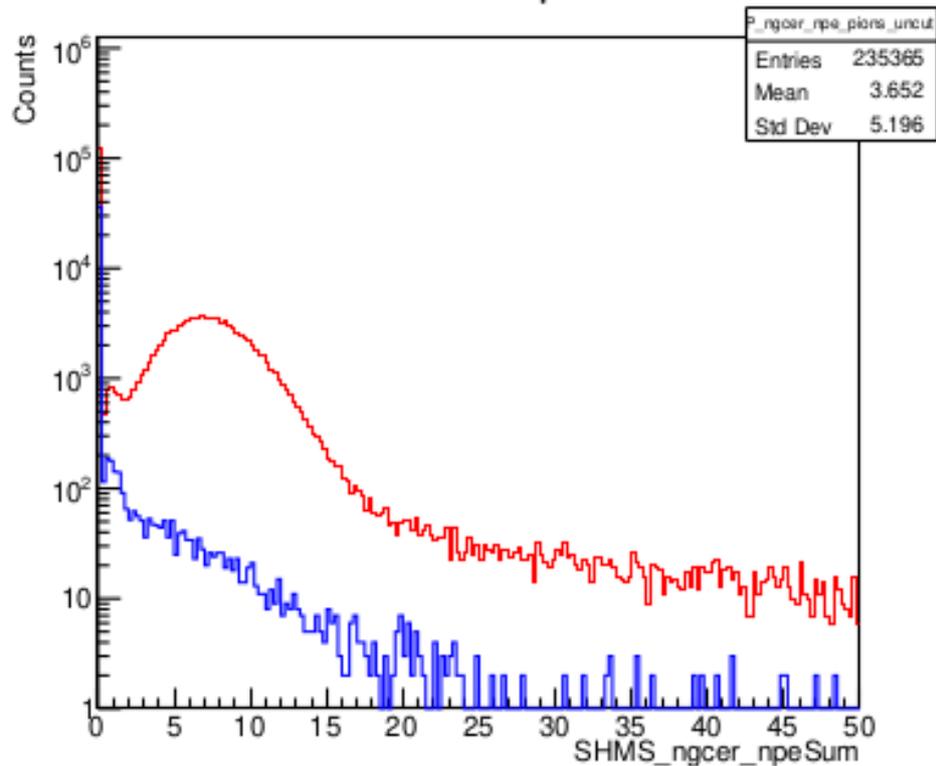


Online
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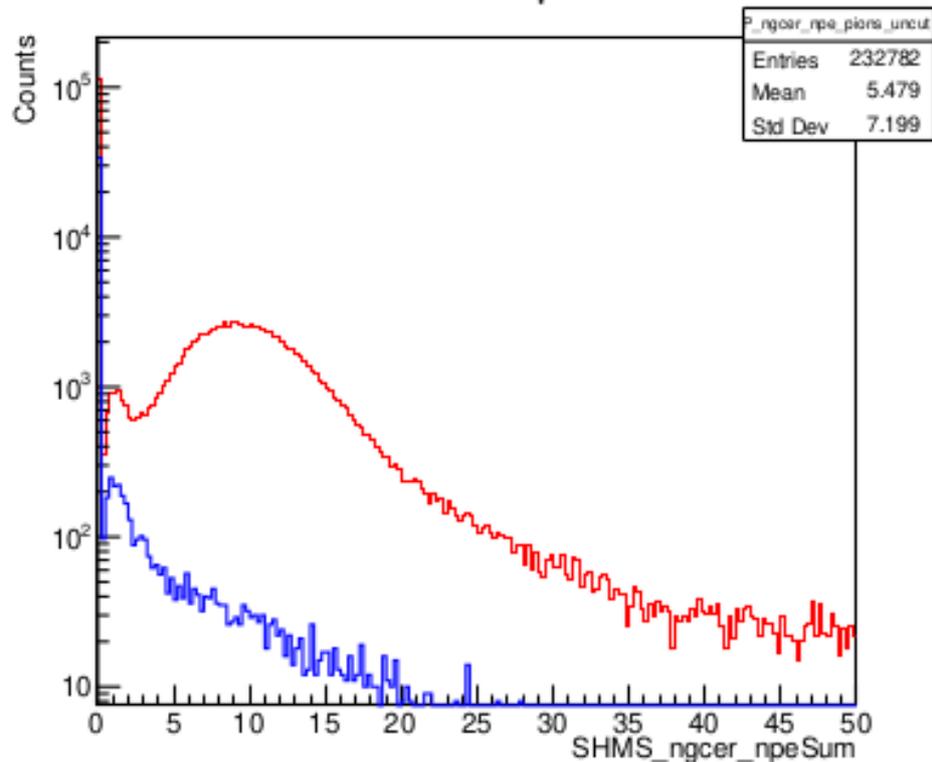


Q^2 6.00 W 2.40 center high ε LD-

Offline
SHMS NGC npeSum



Online
SHMS NGC npeSum



Conclusions

- Calibrated data looks good, by eye a pion cut of around < 3 NPE should work for the entire run period.
- A full PID cut study will have to be done, but that is for later.
- I have begun looking at the HGC calibration, hopefully that can be finished soon.