

Triple-GEM based Polarimeter Tracker for the High- Q^2 Proton Form Factor Measurement (GEp-V) at Jefferson Lab

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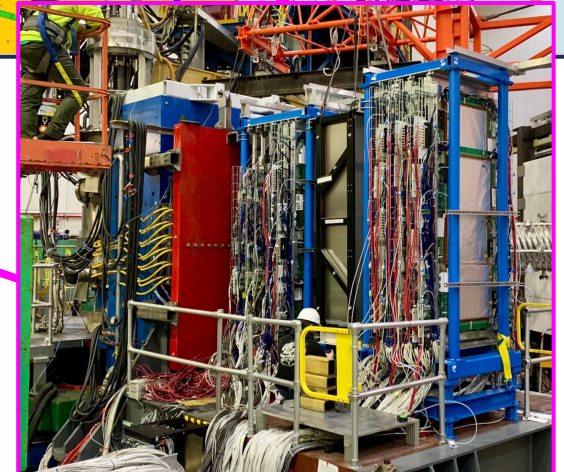
October 2025



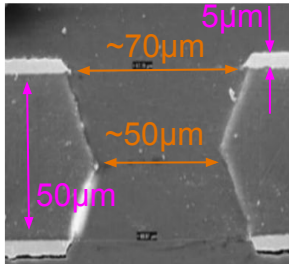
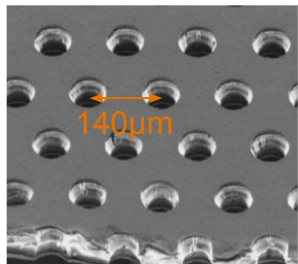
- Super Bigbite Spectrometer(SBS) GEp-V experiment and motivation for GEM based tracking detectors.
- Introduction to Triple-GEM based tracking detectors.
- Triple-GEM based polarimeter tracker for GEp-V.
- Detector Performance during the experiment.
- Recoil Proton Polarimetry and the role of GEM-based tracker.

SBS GEp-V Experiment and Motivation for GEM Tracking

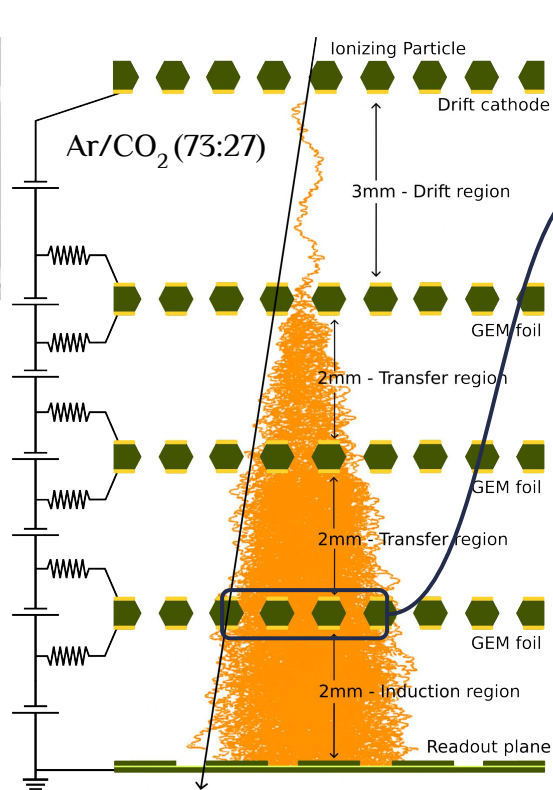
- The SBS GEp-V experiment will provide precision results for proton electromagnetic form factor ratio G_E^p/G_M^p up to 11 GeV^2 .
- Recorded data for two kinematic points.
 - $Q^2 = 5.7 \text{ GeV}^2$ and $Q^2 = 11.1 \text{ GeV}^2$.
- High momentum transfer in the scattering requires the tracking detectors to have.
 - Large solid angle acceptance.
 - High-rate withstanding capabilities up to $\sim 500 \text{ kHz/cm}^2$.
 - Excellent spatial and timing resolution.
- Gas Electron Multiplier (GEM) detectors provide an effective solution.
 - They can handle high rates up to few MHz/cm^2 .
 - Intrinsic spatial resolution $< 70 \mu\text{m}$.
 - Intrinsic timing resolution $< 10 \text{ ns}$.
 - Large area coverage capability.
 - Low thickness minimizing the multiple scattering.



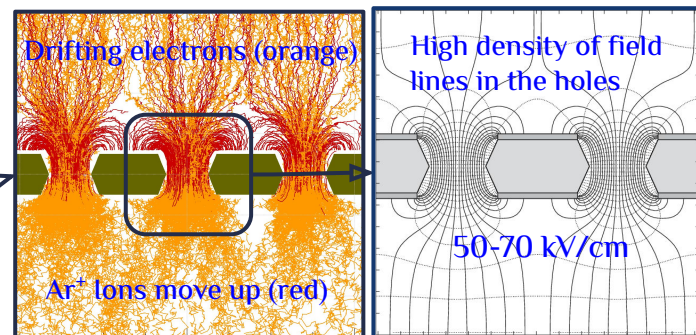
GEM Working Principle and Triple-GEM Detectors



- Invented by physicist Fabio Sauli in 1997 (Sauli, F., NIM A 386 (1997) 531).
- A GEM foil has two copper layers separated by a Polyamide film.
- Chemically etched to produce bi-conical holes.
- High voltage applied between copper layers induces a strong electric field within the hole.

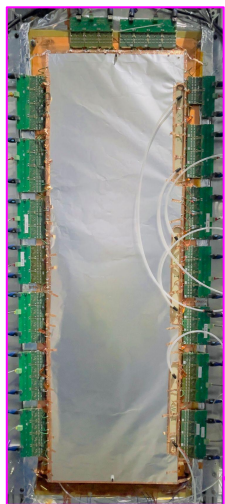


Charge amplification is separated from charge collection → Multi-stage amplification.

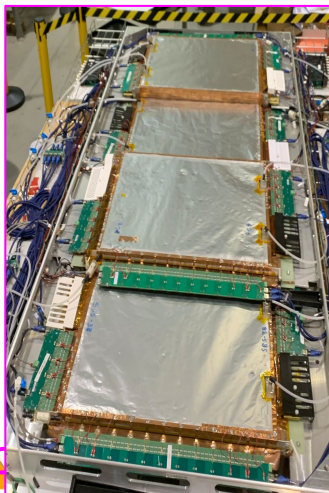
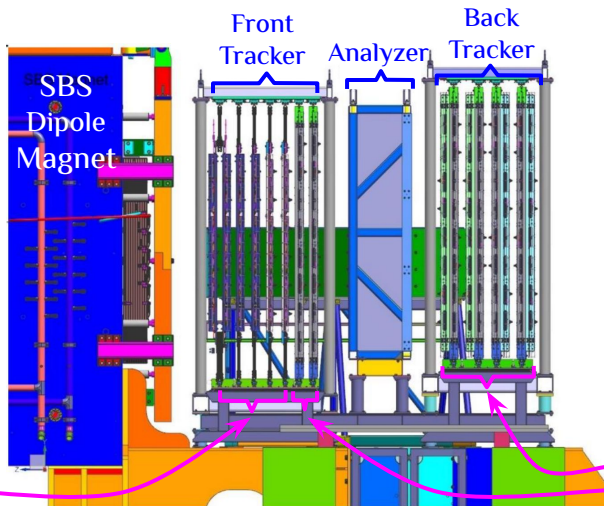


- Incident particle ionizes the gas in the drift region and generates primary electrons.
- Electrons moving through GEM holes are multiplied (about x20 per layer) due to strong electric field.
- An avalanche of electrons moving towards the readout plane induce a signal.
- A 2D strip readout in a GEM detector measures both X and Y coordinates of a particle hit.

Triple-GEM Based Polarimeter Tracker

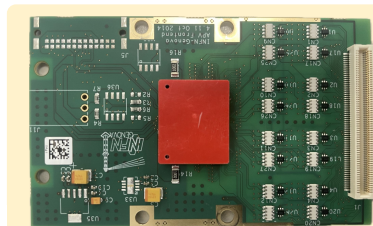


UV/XW layer
40 x 150 sq.cm
Single Modules

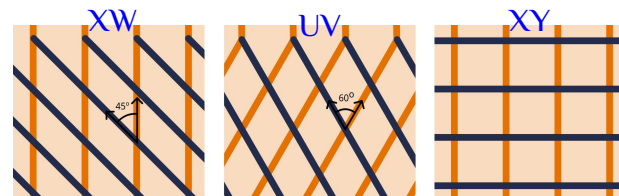
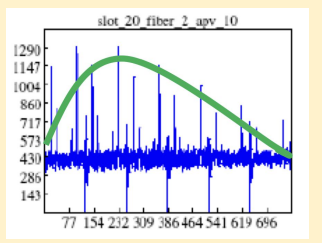


XY layer
Four Modules of
60 x 50 sq.cm each

- 46 GEM modules in Total - Designed and fabricated at UVA.
- Front Tracker:
 - 6 GEM layers of active area 40 x 150 sq.cm.
 - 2 GEM layers of active area 60 x 200 sq.cm.
- Back Tracker (Focal Plane Polarimeter)
 - 8 GEM layers of active area 60 x 200 sq.cm.
- Optimised shorter readout strip geometries in single modules to minimize signal degradation.



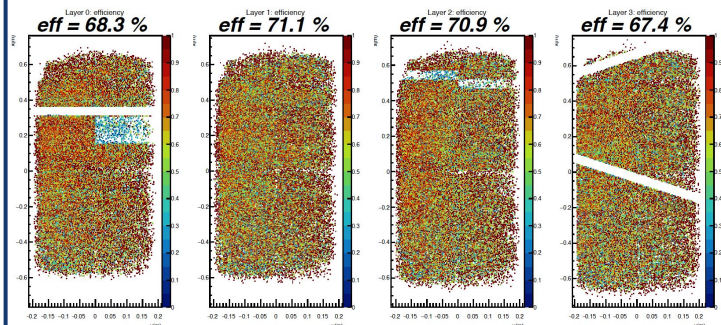
- APV25 based front-end readout electronics.
- 6 times samples.
- 150ns sampling window.



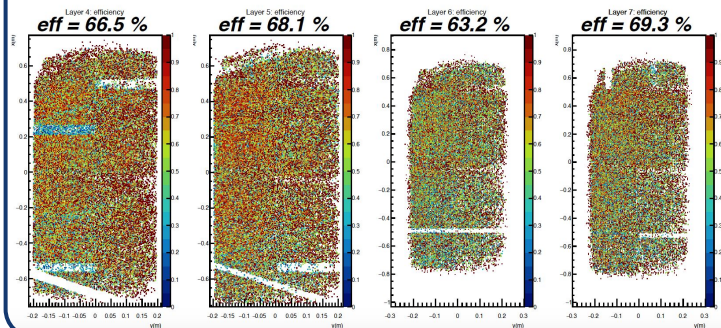
Readout Strip Orientations

GEM Detector Performance during SBS GEp-V Experiment

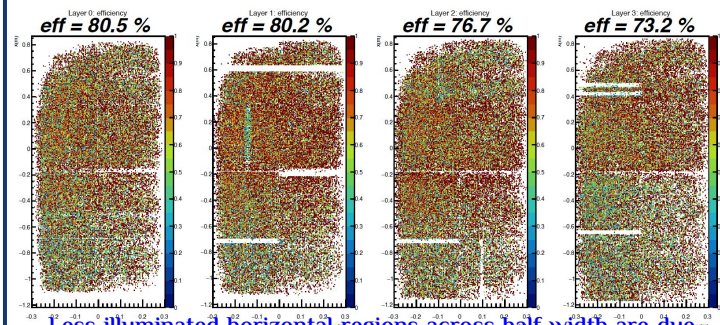
Front Tracker efficiencies ($Q^2 = 11.1 \text{ GeV}^2$)



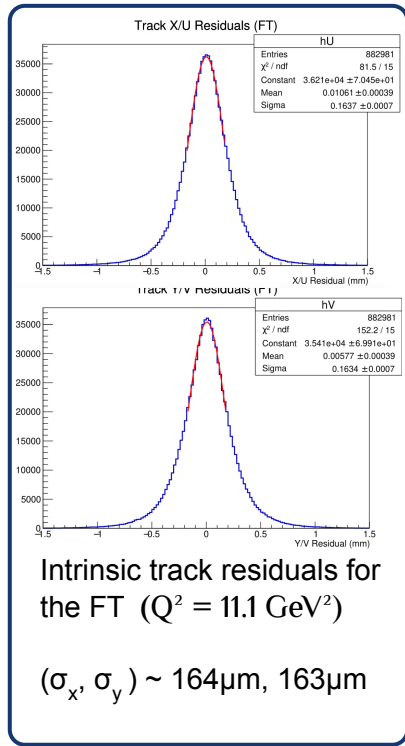
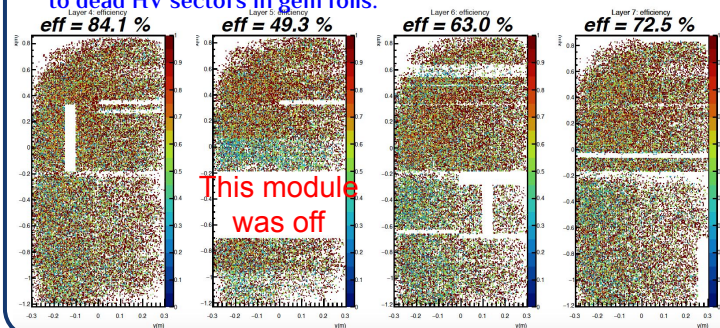
Empty regions are due to readout electronic malfunctions.



Back Tracker efficiencies ($Q^2 = 11.1 \text{ GeV}^2$)



Less illuminated horizontal regions across half width are due to dead HV sectors in gem foils.

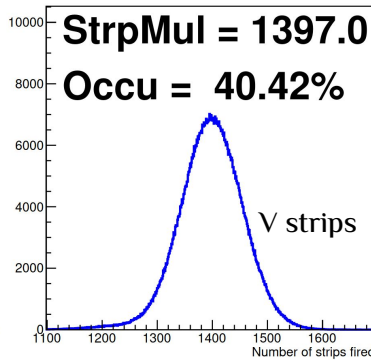
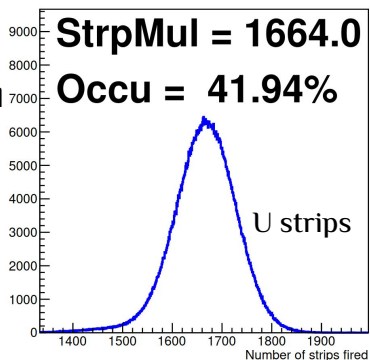


- Plots above show the first look into to the track-based efficiencies of the both front and back trackers.
- Proton tracks used here have passed a series of coincident ECal and HCal timing cuts, GEM track quality cuts, and calorimeter energy and position matching cuts, in order to select elastic ep events.

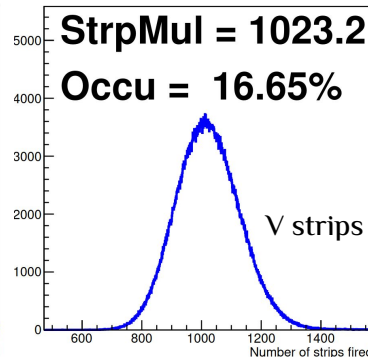
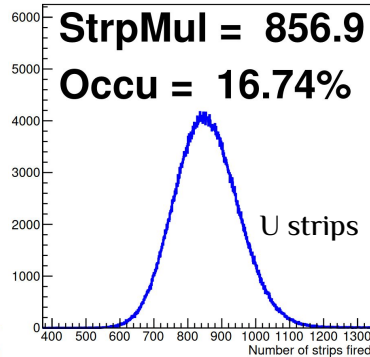
Tracking is a Challenge

- High raw strip occupancy in FT due to soft photon background of rate ~ 500 kHz/cm²
- Partially overlapping background with primary hits make tracking harder.

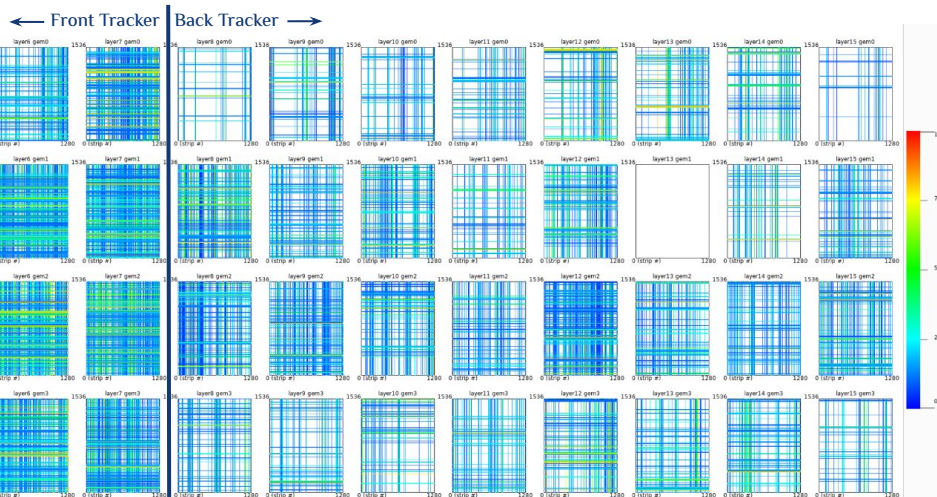
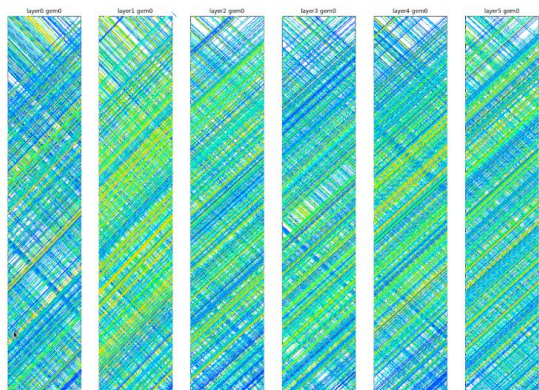
Raw strip occupancies of FT Layer 1



Raw strip occupancies of FPP Layer 1

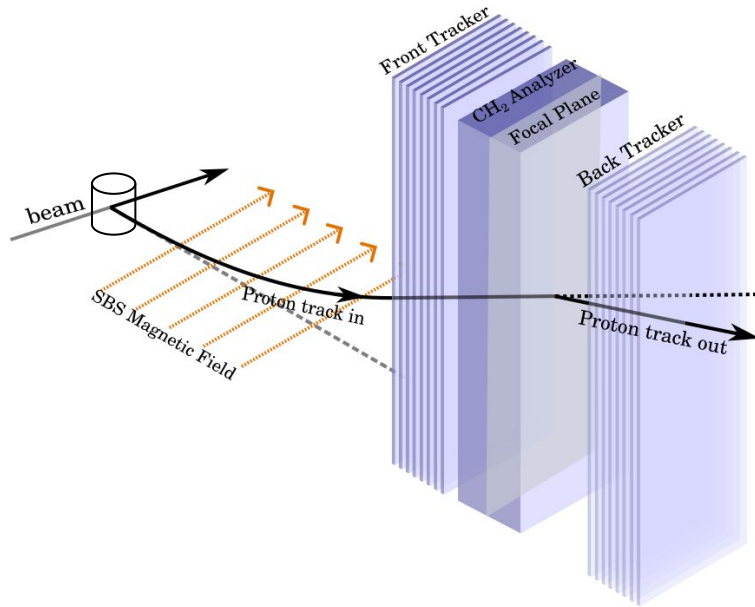


Illuminated lines are strips fired during single 150ns sample window.



Anu will talk about tracking in detail!

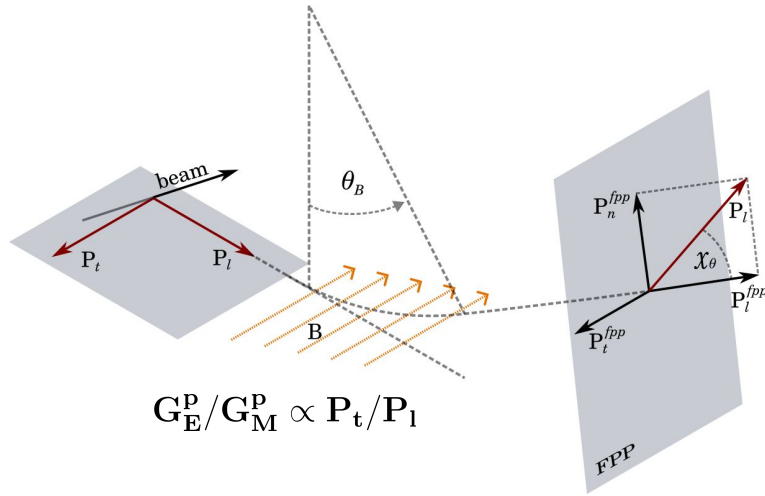
Why do we need two separate precision GEM Trackers?



- A longitudinally polarized electron beam scattered off an unpolarized liquid Hydrogen target.
- Eight GEM layers in the Front Tracker:
 - Reconstruct the proton's momentum after the deflection in the SBS magnetic field.
 - Front tracking defines the incident trajectory at the Focal Plane.
- Secondary scattering in CH₂ analyzer.
- Eight GEM layers in the Back Tracker (FPP):
 - Measure the proton's trajectory after secondary scattering in the CH₂ analyzer.
- Front and back tracks define polar and azimuthal angles of the secondary scattering.

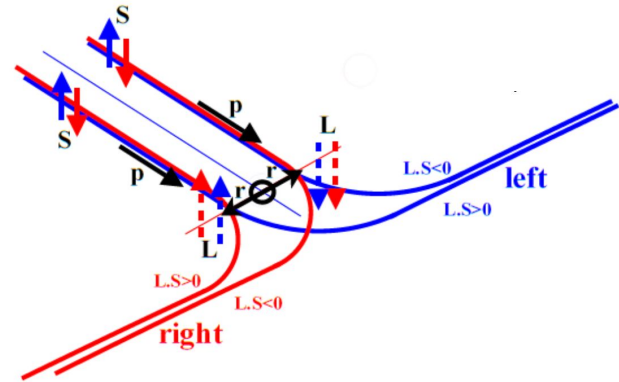
Recoil Proton Polarimetry

- Spin-polarized ensemble of electrons impart net polarization to the scattered ensemble of protons.

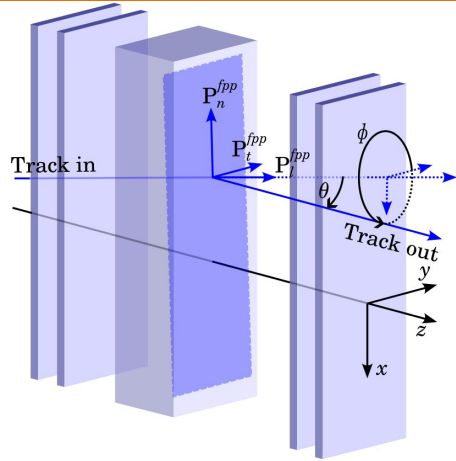


- Polarization component can be measured through a suitable secondary scattering.
 - The Recoil Proton Polarimetry is based on the spin-orbit coupling of the proton-nucleus in secondary scattering.
- (Akhiezer & Rekalov, *Dokl. Akad. Nauk SSSR* 180 (1968))

- The spin orbit force has the form $\mathbf{L} \cdot \mathbf{S}$. Since $\mathbf{L} \cdot \mathbf{p} = (\mathbf{r} \times \mathbf{p}) \cdot \mathbf{p} = 0$, spin-orbit force is insensitive to P_l .
- Spin precession in the magnetic field of SBS rotates P_l to a transverse component.
- The spin-orbit force preferably deflects the proton in the direction $\mathbf{p} \times \mathbf{S}$ (parallelly or anti-parallelly based on the sign of $\mathbf{L} \cdot \mathbf{S}$).



Recoil Proton Polarimetry Continued...



- If one spin state is dominant, this creates an asymmetry in the azimuthal angular distribution of the secondary scattering.

$$f^{\pm}(\theta, \phi) = \frac{\epsilon(\theta, \phi)}{2\pi} \left[1 \pm A_y \left(P_t^{fpp} \sin \phi - P_n^{fpp} \cos \phi \right) \right]$$

For negative and positive helicity states

Instrumental asymmetry that describes non-uniformities in detector response

Carbon analyzing power

- Fit the distribution to obtain polarization components simultaneously.

$$A = \frac{f^+ - f^-}{f^+ + f^-} = A_y \left(P_t^{fpp} \sin \phi - P_n^{fpp} \cos \phi \right)$$

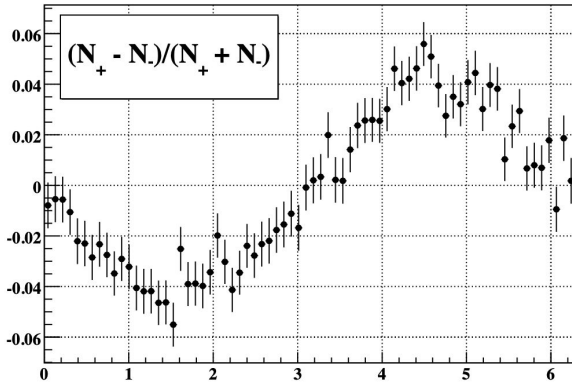
$$\mu_p \frac{G_E^p}{G_M^p} = -\mu_p \frac{E_e + E'_e}{2M_p} \tan \frac{\theta_e}{2} \left(\frac{P_t^{fpp}}{P_n^{fpp}} \sin \chi_\theta + \gamma_p (\mu_p - 1) \Delta\phi \right)$$

Precession angle

Fringe field correction

μ_p — proton magnetic moment, E_e — incident beam electron energy

E'_e — scattered electron energy, θ_e — electron scattering angle, M_p — proton mass



Example from GEp-III (Puckett, A. J. R., arXiv:1508.01456 (2015)).

Overview and Upcoming Work

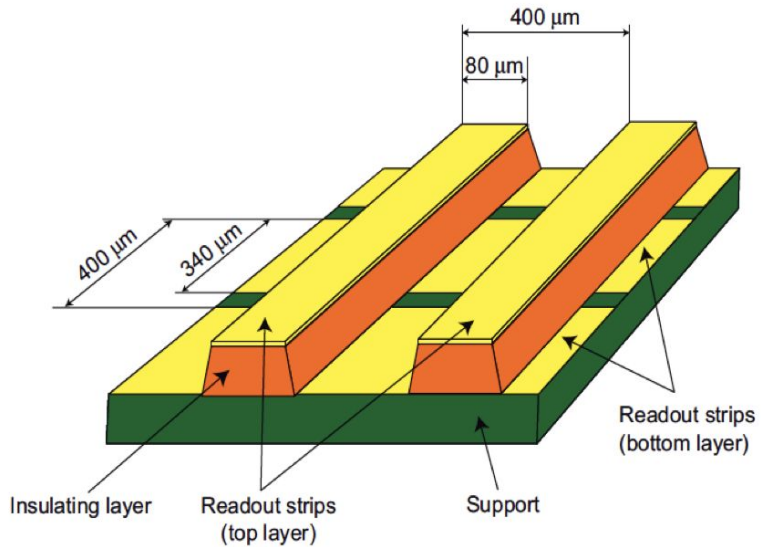
- SBS GEp-V experiment uses the Recoil Polarization technique to obtain the proton electromagnetic form factor ratio G_E^p/G_M^p .
- The Triple GEM-based polarimeter tracker demonstrated reliable and stable performance even at unprecedented rates.
- Initial tracking efficiencies: FT layers 60%-70%, FPP layers 70%-80%; improvements anticipated with optimized reconstruction.
- Ongoing work on GEM detector calibrations and tracking reconstruction improvements.

Acknowledgements

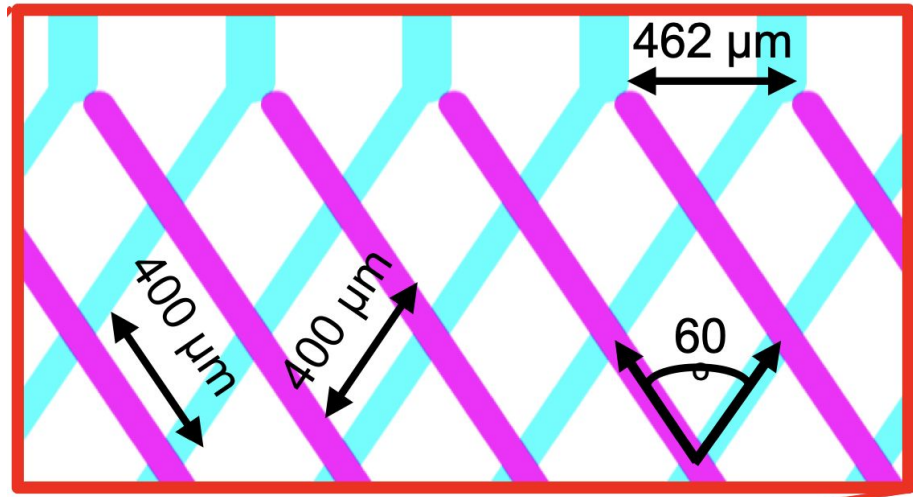
- The Super Bigbite Collaboration and SBS GEp-V spokespeople: Andrew Puckett, Bogdan Wojtsekhowski, Evaristo Cisbani, Lubomir Pentchev, Mark Jones, Nilanga Liyanage.
- GEp-V Graduate Students : Kip Hunt, Mahmoud Gomina, Ben Spaude, Jhih-Ying Su, Jacob McMurtry.
- UVa Gas Detector Group: Prof. Nilanga Liyanage, Research Assistant Prof. Huong Thi Nguyen, Dr. Asar Ahmed, Vimukthi Gamage, Bhasitha Dharmasena, Jacob McMurtry, Nithya Kularatne.
- US Department of Energy, Office of Science, Office of Nuclear physics award number DE-FG02-03ER41240

Thank You!

Backup Slides



COMPASS X-Y strip readout



Avalanche multiplication occurs in all gases, but effective operation needs a specific *gas mixture*.

Desired properties of a good mixture:

- Low working voltage (low ionization potential)
- Stable high-gain operation
- High rate capability (fast recovery)
- Good proportionality

Noble gases (e.g., Ar, He, Ne, Kr, Xe) are the are a good choice because they have no molecular structure. So there is no energy absorption in collisions.

Argon produces more primary ionization than He or Ne; Kr and Xe are even better but expensive.

But, pure argon is unstable at high gain due to:

- High excitation energy (11.6 eV) → emission of energetic UV photons.
- UV photons cause photoemission from the cathode → secondary avalanches.
- Leads to self-sustaining discharge (breakdown).

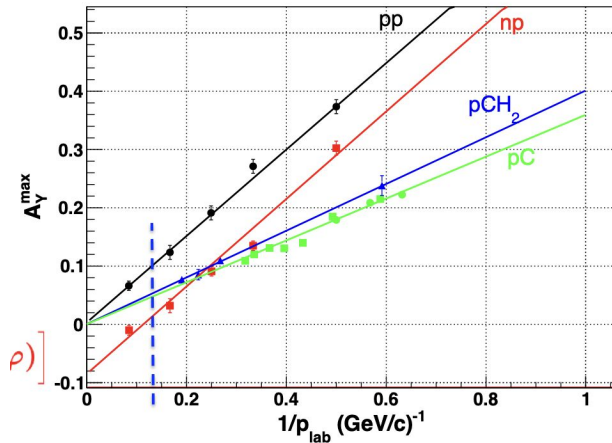
Adding polyatomic gases (e.g. CH₄, isobutane, CO₂) improves stability because they have many non-radiative vibrational and rotational modes that absorb energy.

- In the Ideal Dipole Approximation:

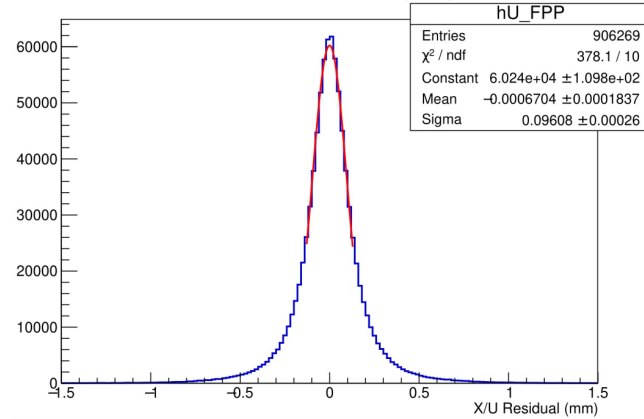
$$\chi_\theta \equiv \gamma \kappa_p \theta_{\text{bend}}$$

κ_p - Proton's anomalous magnetic moment

A_Y analyzing power vs.
inverse proton momentum



Track X/U Residuals (FPP)



Track Y/V Residuals (FPP)

